



We can write for $B \neq \phi$

$$\cos \psi(t) = \cos(\psi_0 + \Omega t)$$

with -

$$\left. \begin{aligned} \text{lat} &= \lambda \\ \text{lon} &= \psi \end{aligned} \right\}$$

$$V_x(t) = \Omega R \cos(\lambda) \cdot \cos \psi(t) = \Omega dx(t)$$

so

$$dx(t) = \int_{\psi_0}^{\psi_0 + \Omega t} R \cos(\lambda) \cos \psi(t) dt$$

$$dx(t) = R \cos(\lambda) \int_{\psi_0}^{\psi_0 + \Omega t} \cos \psi(t) dt$$



$$dx(t) = R \cos(\lambda) [\sin(\psi_0 + \Omega t) - \sin(\psi_0)]$$

Solar differential rotation

$$\Omega_{(lat)} = a + b \sin^2(lat) + c \sin^4(lat)$$

$$V_y(t) = \Omega R \cos(\lambda) \sin \psi \sin(B \neq \phi) = \Omega dy(t)$$

$$dy(t) = \int R \cos(\lambda) \sin \psi \sin(B \neq \phi) dt$$

$$dy(t) = R \cos(\lambda) \sin(B \neq \phi) \int_{\psi_0}^{\psi_0 + \Omega t} \sin \psi(t) dt$$

$$dy(t) = R \cos(\lambda) \sin(B \neq \phi) [-\cos(\psi_0 + \Omega t) + \cos(\psi_0)]$$

$$dy(t) = -R \cos \lambda \sin B \neq \phi [\cos(\psi_0 + \Omega t) - \cos(\psi_0)]$$

Conclusion

To deconvolve the solar image for the CST, we used $dx(t)$ and $dy(t)$ formulae